Auroral Precipitation Models: A Potential Future Topic for GGCM Metrics and Validation

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All of you - the Audience

Existing Models

- Ovation model Newell et al.

CCMC is collaborating with AFWA in validating both of them (Hardy and Ovation)).

- Ovation Prime Newell et al. (running at CCMC, runon-request soon) JGR 2009; 2010
- Zhang and Paxton Auroral Model
- Combining a Global MHD and an inner magnetosphere model to determine the boundaries (one already running at CCMC)

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Data for Validation

- DMSP particle data (not for the models which are based on the DMSP data)
- Auroral imaging data

IMAGE/FUV, May 2000-Dec, 2005 <u>Stephen Mende/Harald Frey</u>
POLAR/UVI, March, 1996-1999, 2007 <u>Kan Liou</u>
DMSP/SUSSI, 2005 – present <u>Larry Paxton/Yongliang Zhang</u>
TIMED/GUVI Feb. 2002 –Nov 2007 <u>Larry Paxton/Yongliang Zhang</u>

Other data

Particle data (local, better determination of boundaries) versus Imaging data (global, but compromises in boundary determination?) need both!

Potential events for contemplation

- One Quiet time
- One during Steady Magnetospheric Convection (SMC) period
- One during active substorm period
- One during superstorm time

Working with experts/you in selecting events

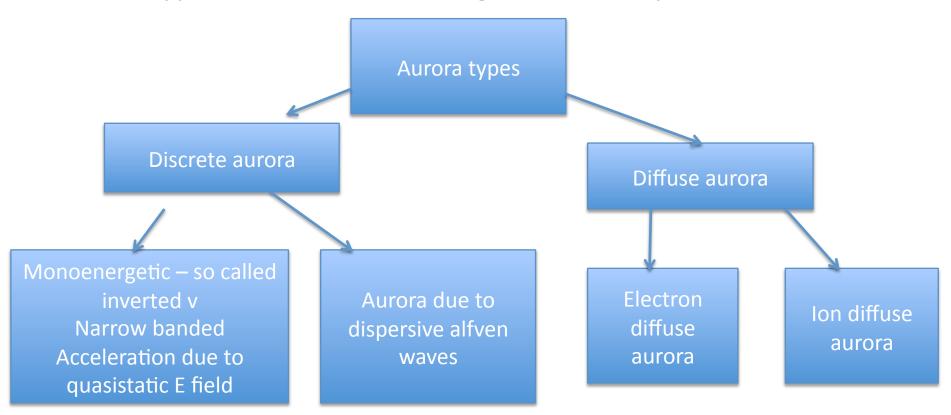
What to evaluate exactly

- Correlation coeff? Morphology?
- Characteristic energy (average energy) and energy flux of precipitating electrons – from those two, it is also easier to convert to conductance, which is very important to the dynamics of M-I coupling.
- Boundaries equatorward and poleward boundaries, their definition needs to be agreed upon

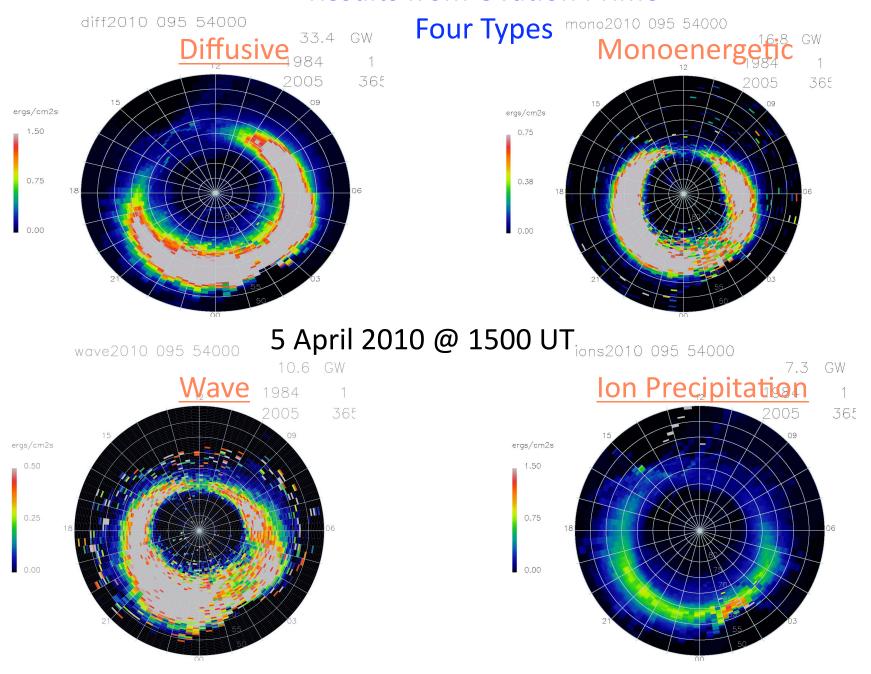
END

Ovation Prime

- Not just higher resolution than Ovation
- Parameterized by solar wind coupling functions, not just Kp
- Four types of aurora, including seasonal dependence

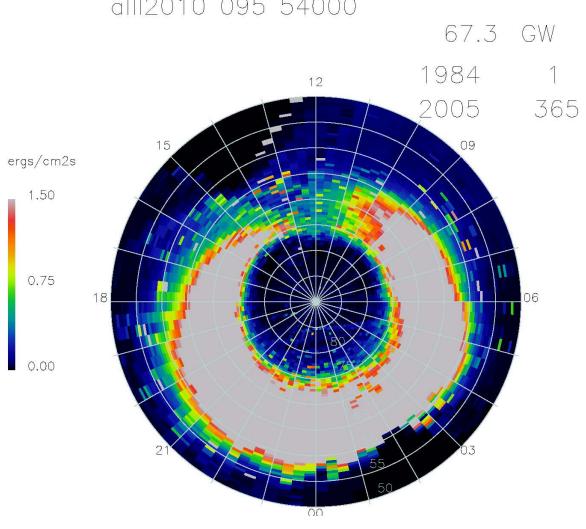


Results from Ovation Prime



Results from Ovation Prime Sum of all four types





Main References

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