

2016 Summer GEM
Mid-tail (60 Re) modeling challenge
“Tail Environment and Dynamics at Lunar Distances” &
“Modeling Methods and Validation” FGs

ARTEMIS observations for a northward IMF event:

IMF remains northward for 48 hr from 13 Feb to 14 Feb 2014

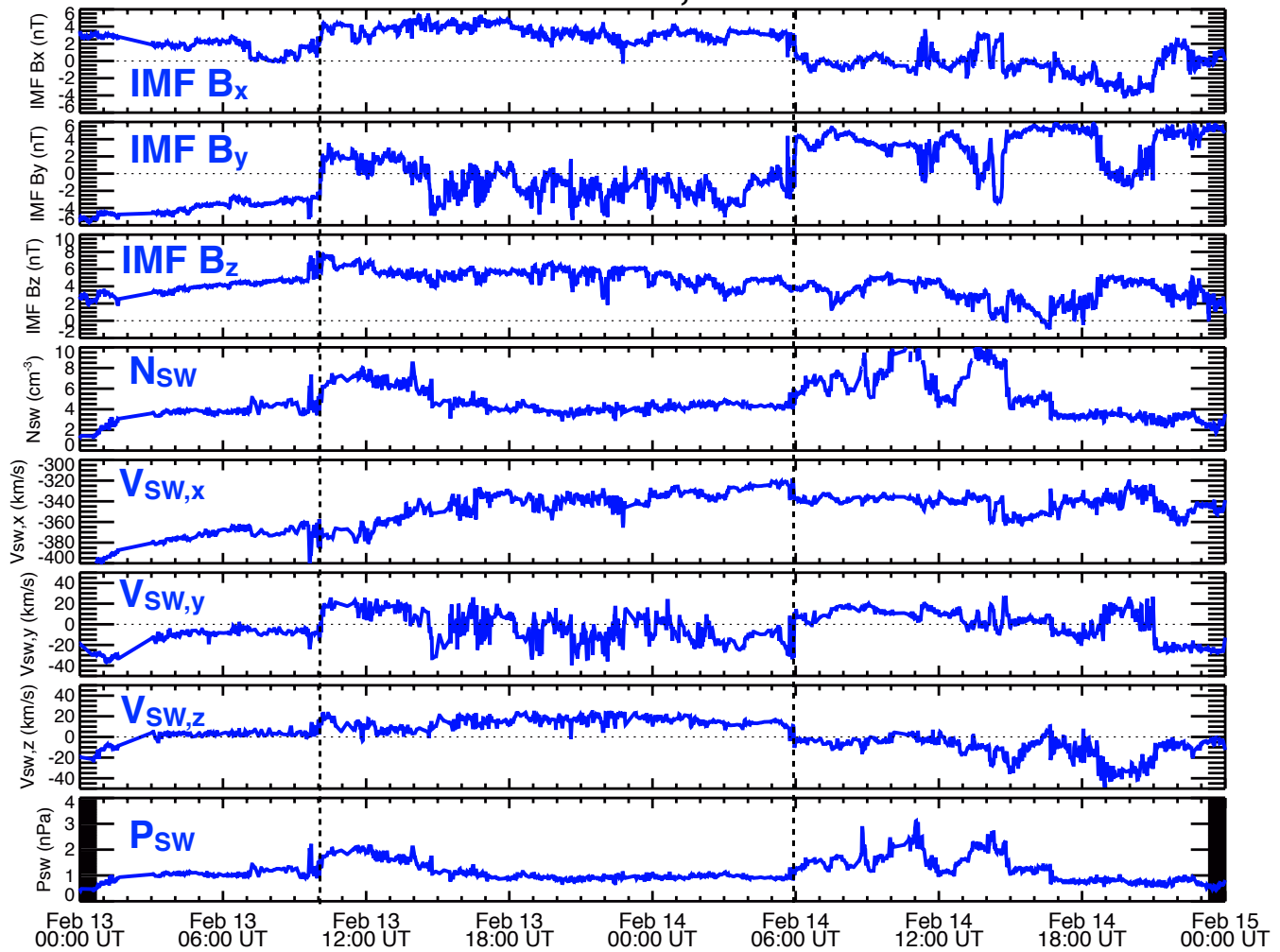
The two ARTEMIS probes at $X = -60 R_E$ moving from $Y = 23$ to $3 R_E$ and are close to the plasma sheet.

The two probes (separation ~ 1 to $3 R_E$) provide determination of spatial and temporal variations.

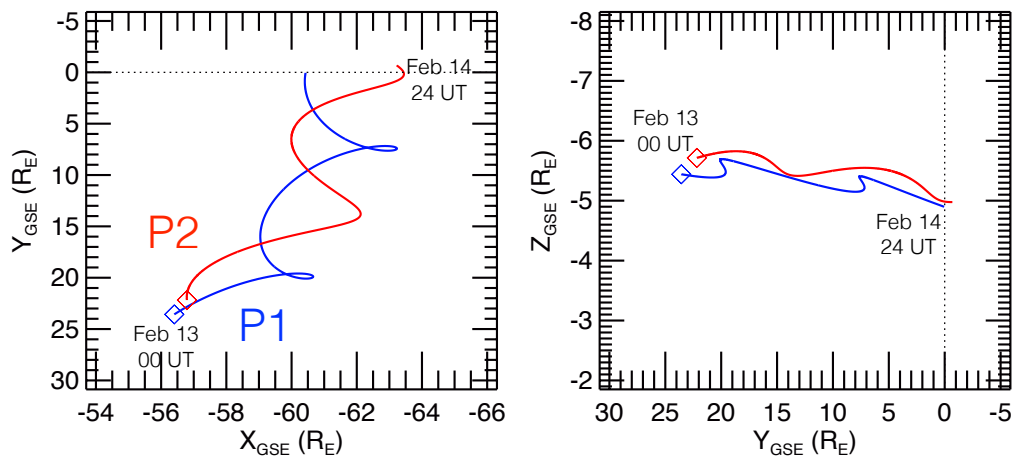
We would like to investigate the ability of global models (both MHD and hybrid models) on the following questions:

1. How does the mid-tail configuration response to changes in IMF B_y and solar wind dynamic pressure?
2. What processes cause the 10 to 30 min perturbations in the mid-tail plasma sheet?
3. What are characteristics of plasma flows in mid-tail plasma sheet?

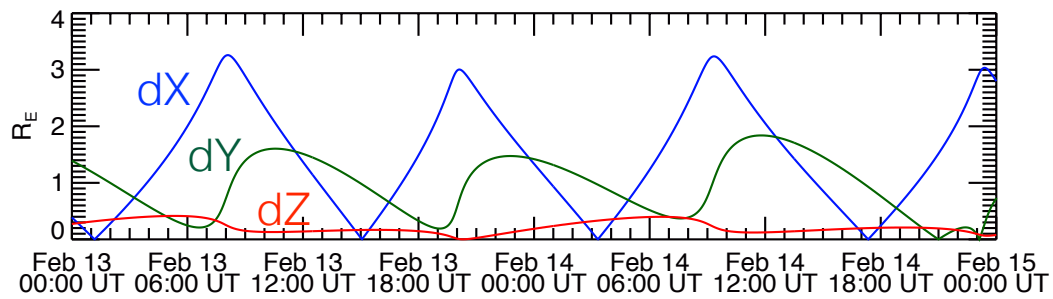
SW/IMF conditions, 13-14 Feb 2014



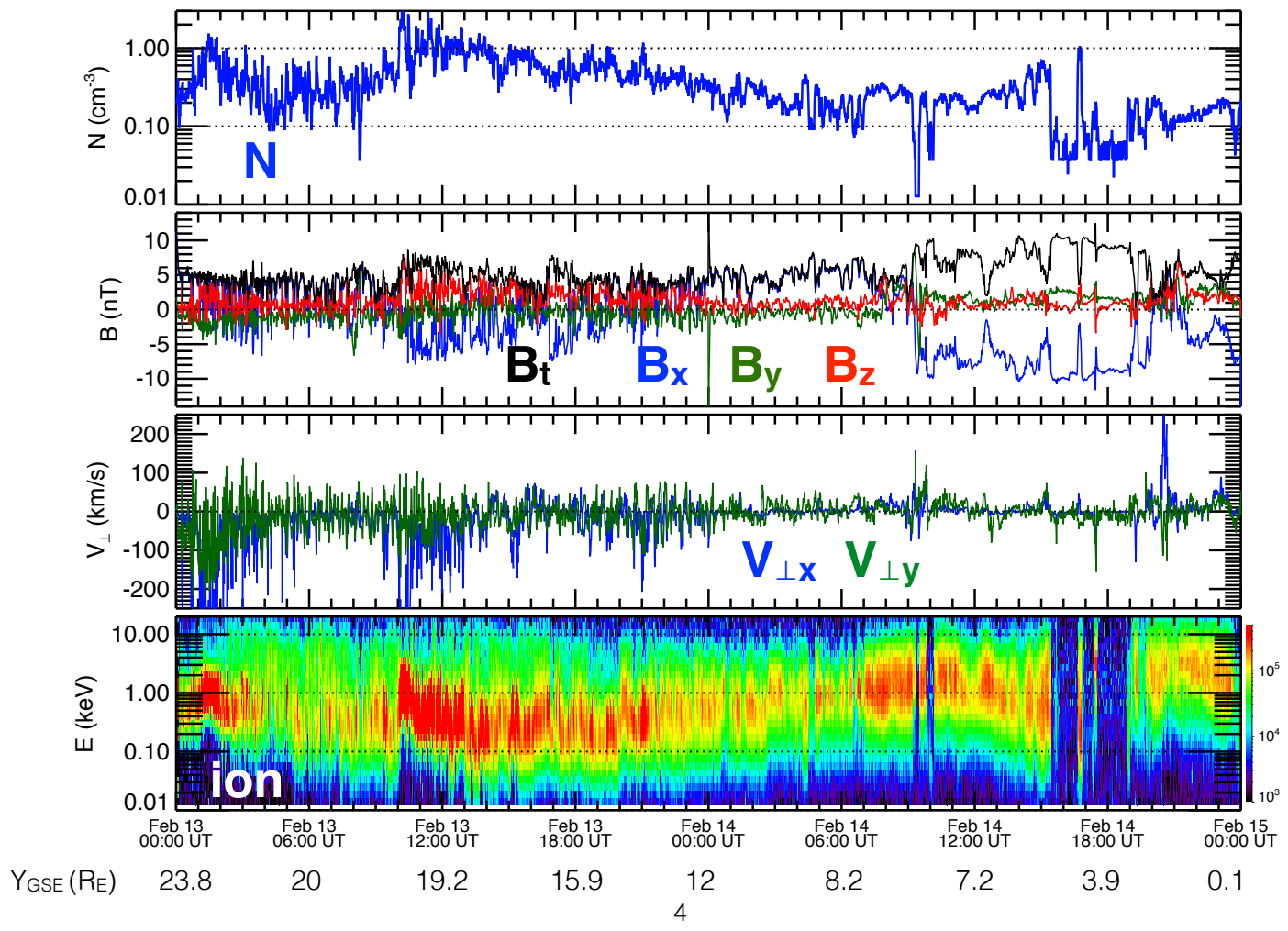
locations of ARTEMIS P1 and P2



Separation between P1 and P2



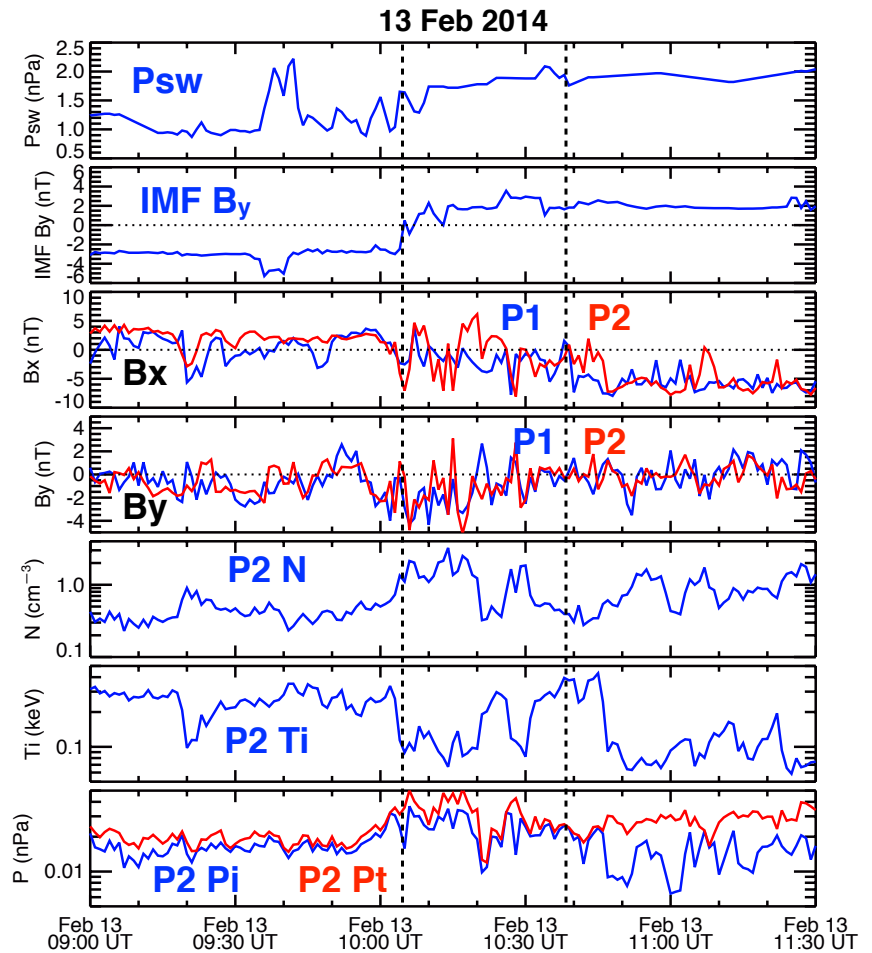
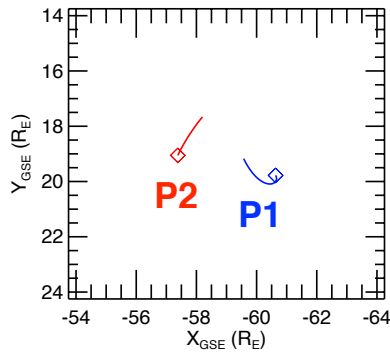
ARTEMIS P1 observation



1. How do mid-tail respond to IMF By and Psw changes?

P1 and P2 responses to the 1st change

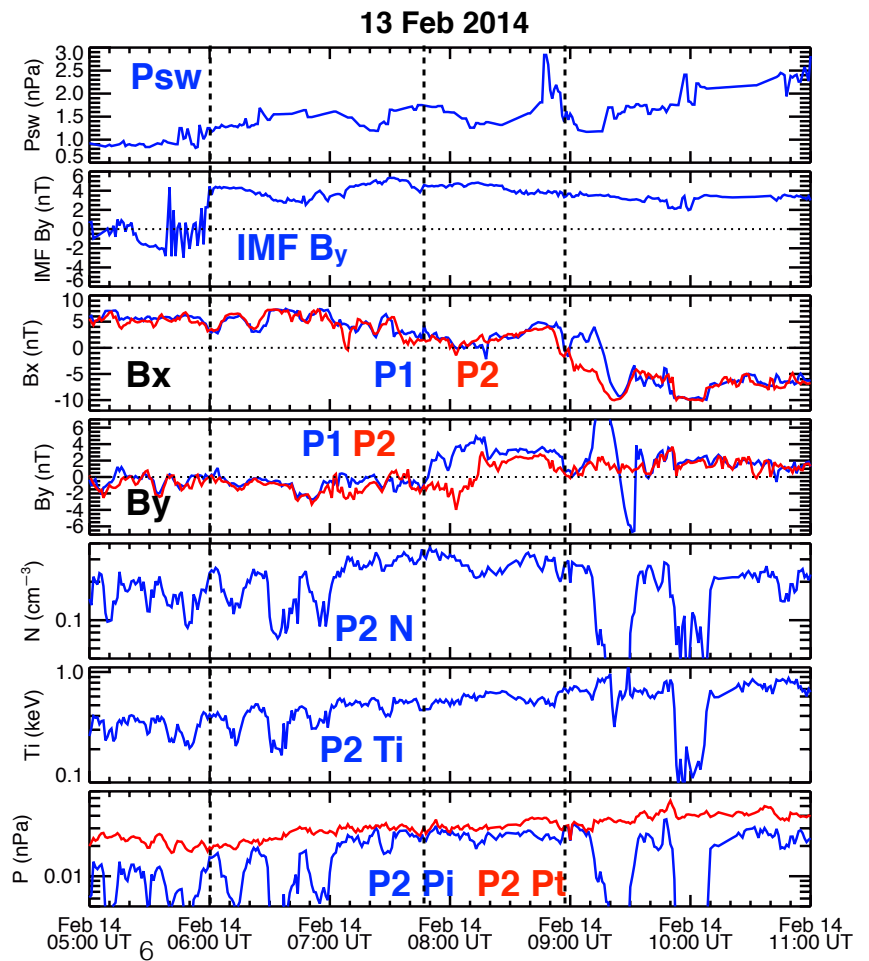
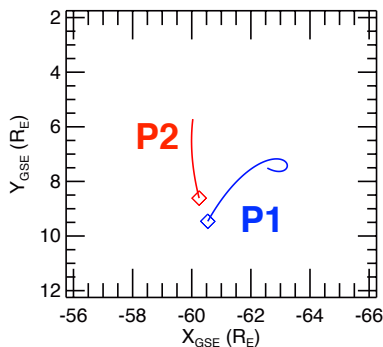
- P1 P2 moved from N hemisphere to S hemisphere ~30 min after IMF By direction change
- In response to the P_{sw} increase from ~1 to 2 nPa, N increases from 0.3 to 0.9 cm^{-3} but Ti decreases from 0.3 to 0.1 keV, Pi increases slightly from 0.016 to 0.02 nPa



1. How do mid-tail respond to IMF By and Psw changes?

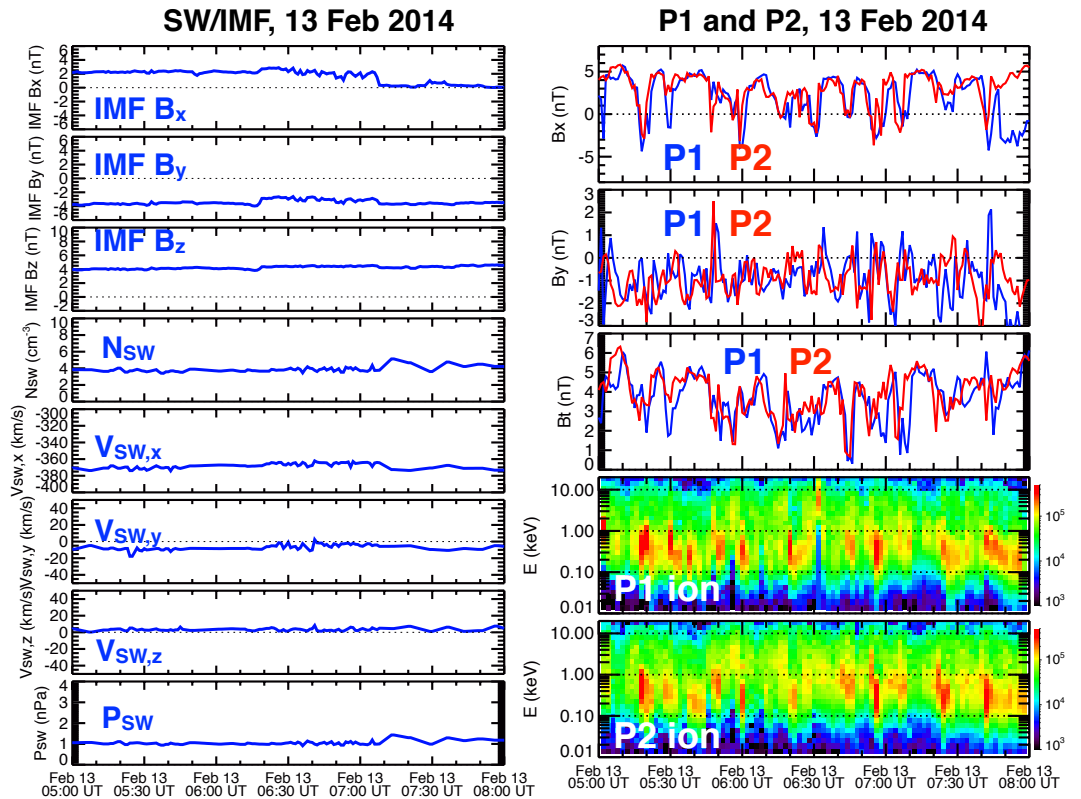
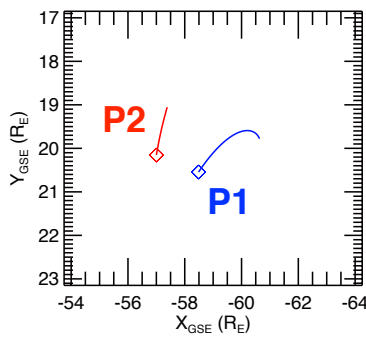
P1 and P2 responses to the 2nd changes

- P1 P2 moved from N hemisphere to S hemisphere ~3 hr after IMF By direction change. **Why did it take much longer than the 1st one? Because of smaller |Y|?**
- In response to gradual Psw increase from ~0.9 to 1.5 nPa before 08 UT, N increases from 0.17 to 0.22 cm⁻³ and Ti increases from 0.4 to 0.6 keV, Pi increases from 0.017 to 0.03 nPa



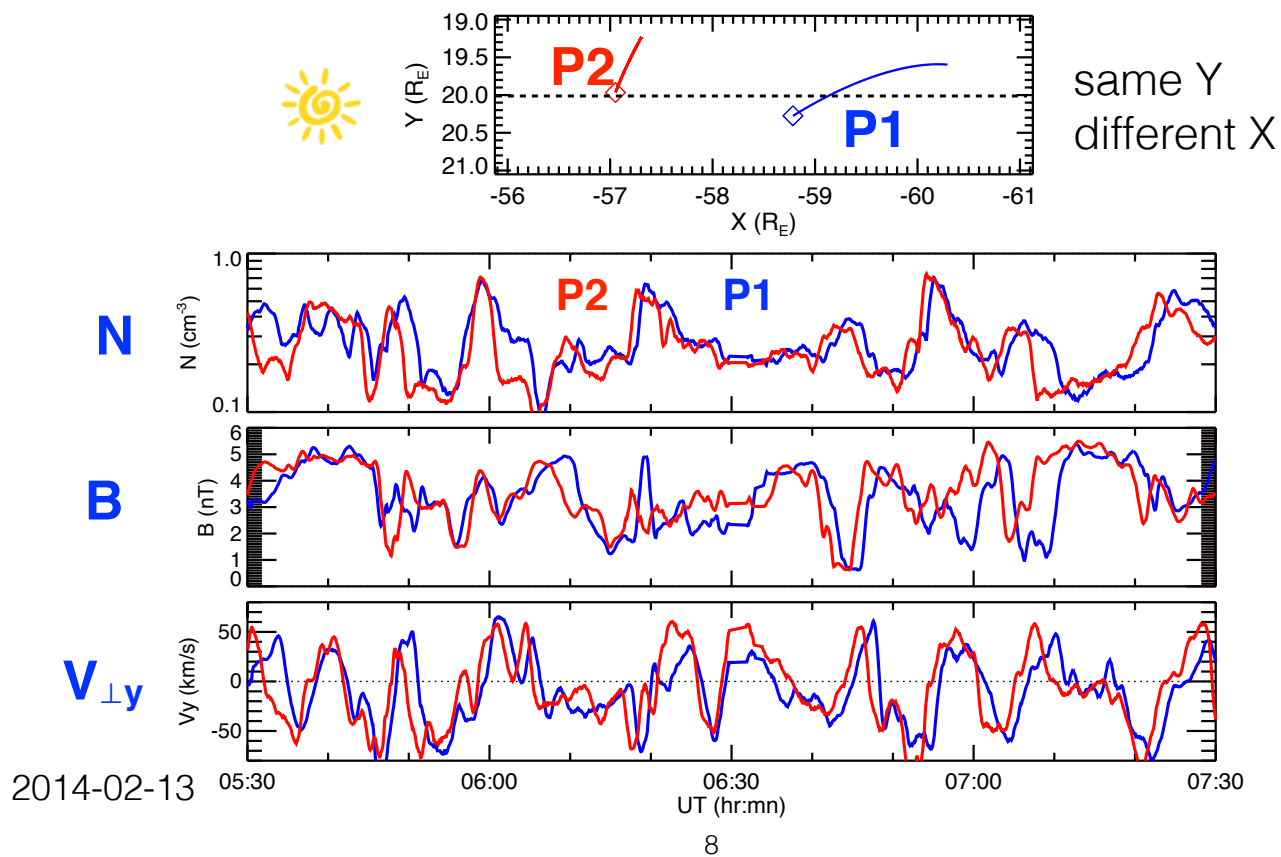
2. What processes cause the 10-30 min perturbations in the mid-tail plasma sheet during a period when SW/IMF are steady?

- From ~04-09 UT on 13 Feb all SW/IMF parameters are relatively steady
- quasi-periodic fluctuations in magnetic field and plasma with appearance of cold plasma.



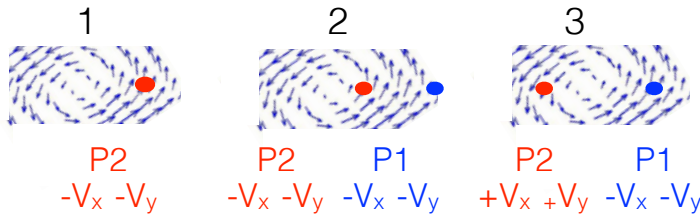
2. What processes cause the 10-30 min perturbations in the mid-tail plasma sheet during a period when SW/IMF are steady?

- P2 closer to the Earth and saw perturbations first, indicating the perturbations propagate tailward



2. What processes cause the 10-30 min perturbations in the mid-tail plasma sheet during a period when SW/IMF are steady?

- Flow direction perturbations consistent with a tailward moving vortex



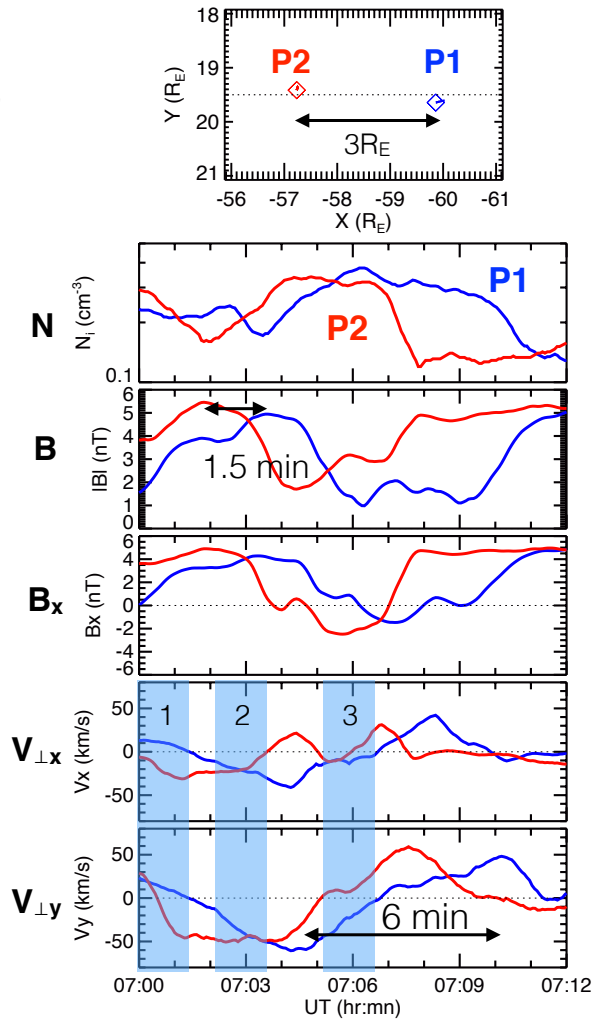
- From the delay time between P1 and P2 (1.5 min) the speed of vortex is

$$V_{\text{vortex}} = 3 R_E / 1.5 \text{ min} = 140 \text{ km/s}$$

- From 6 min between two $|V_y|$ peak, the X-scale of the vortex is

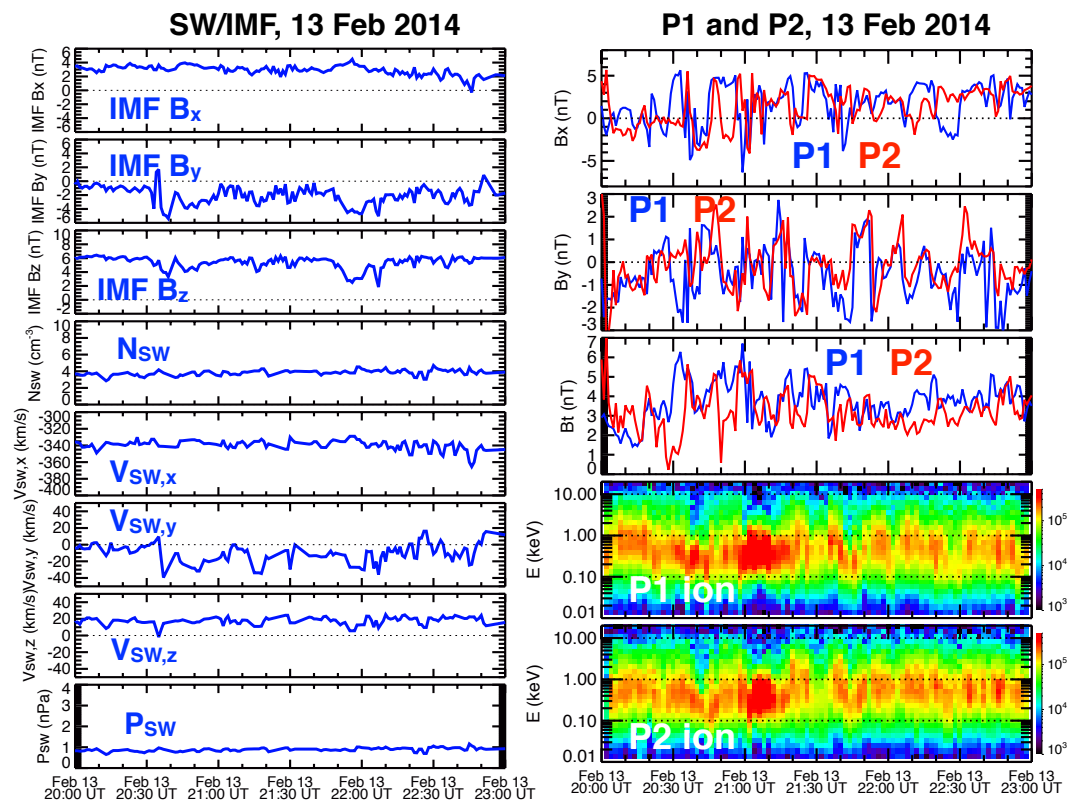
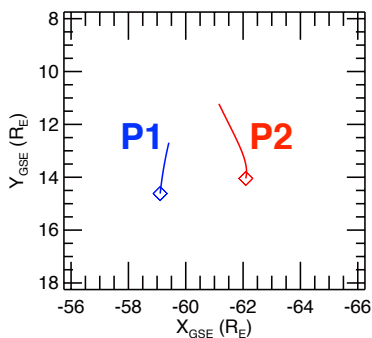
$$X_{\text{vortex}} = V_{\text{vortex}} * 6 \text{ min} = 8 R_E$$

- Are the fluctuations caused by K-H waves?

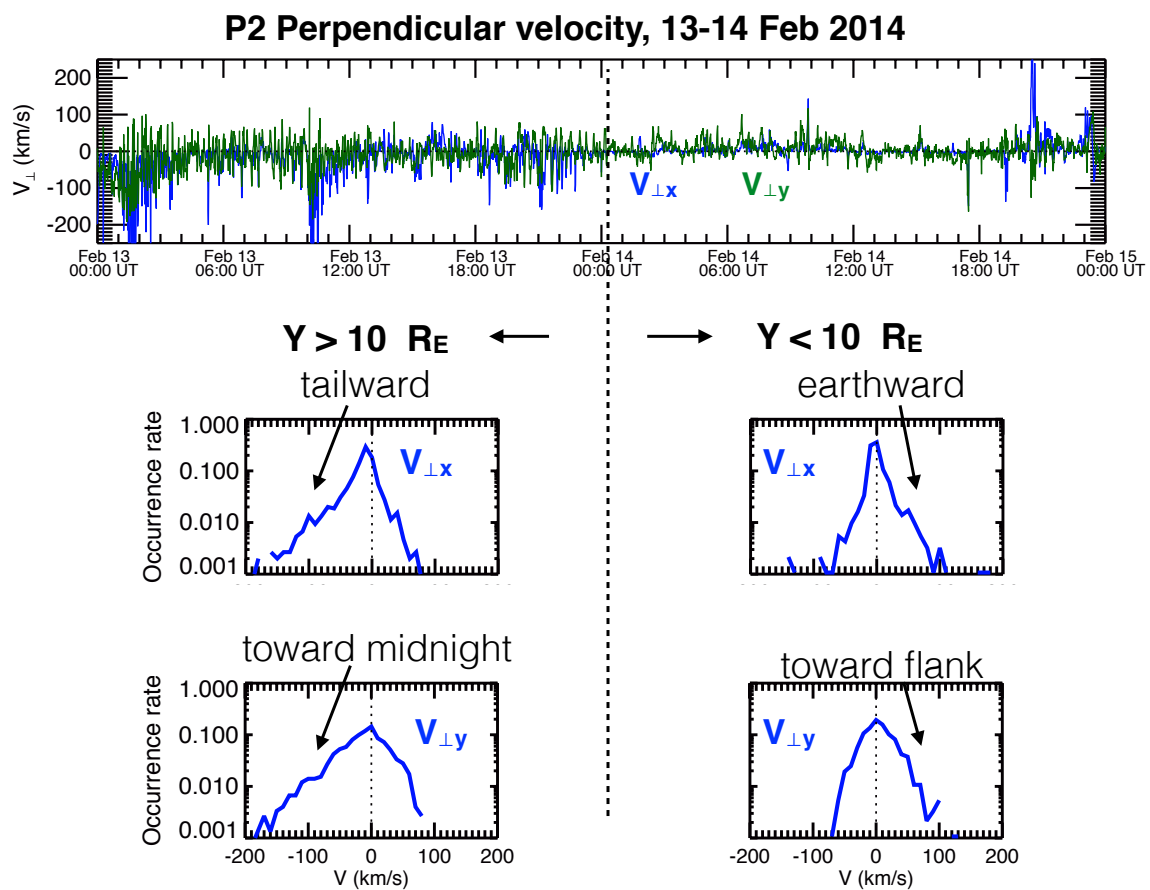


2. What processes cause the 10-30 min perturbations in the mid-tail plasma sheet during a period when there are small SW/IMF perturbations?

- Another period of magnetic field and plasma fluctuations, **more irregular**. Are the fluctuations current sheet flapping motion?
- **Are the fluctuations caused by the small fluctuations in IMF B_y or $V_{sw,y}$?**



3. What are characteristics of plasma flows in mid-tail plasma sheet?

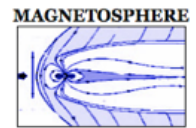


participating global models

1. LFM: Slava Merkin
2. OpenGGCM: Joseph Jensen
3. GUMICS: Ilja Honkonen
4. BATS-R-US: (Daniel Welling)
5. Auburn University hybrid code: Yu Lin, Xueyi Wang



global models on CCMC



- Also run LFM, OpenGGCM, BATS-R-US, and GUMICS on CCMC as baseline

		cells	$\Delta X (R_E)$ at $X = -60 R_E$	$\Delta Y (R_E)$	$\Delta Z (R_E)$
LFM	SM	106x96x128	1.2-1.8	1.2-1.8	1.2-1.8
OpenGGCM midtail emphasis grid	GSE	9M	0.6	0.4-1.2	0.2-0.6
SWMF/ BATS-R-US	GSM	2M	0.5-2	0.5-2	0.5-2