

## Modeling Ionospheric Super-Fountain Effect Based on the Coupled TIMEGCM-SAMI3

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## SAMI3

- A 3d global model of the ionosphere and plasmasphere, extending from 85km to a few R<sub>E</sub>;
- Non-orthogonal, non-uniform, fixed grids, and the spatial resolution decreases with increasing altitude;
- Solving time-dependent equations of continuity and momentum for electrons and 7 ion species (i.e., H<sup>+</sup>, He<sup>+</sup>, O<sup>+</sup>, O<sub>2</sub><sup>+</sup>, N<sup>+</sup>, N<sub>2</sub><sup>+</sup>, and NO<sup>+</sup>);
- Solving temperature/energy equations for 3 ion species (H<sup>+</sup>, He<sup>+</sup>, and O<sup>+</sup>) as well as electrons;
- Including the ion inertia term in the ion momentum equation.

## The Coupled TIMEGCM-SAMI3

- Neutral species (i.e., H, O, O<sub>2</sub>, N, N<sub>2</sub>, and NO), but He from NRLMSISOO;
- Neutral temperature, neutral wind, (and electric field);
- Auroral energy flux and mean energy from AMIE-TIMEGCM;
- Implemented the auroral ionization code in SAMI3;
- The 5-min TIMEGCM outputs are interpolated spatially and temporally to drive SAMI3;
- Currently, it is one-way coupling!

## Distributions of IEF, PCP, and Vertical Ion Drift



[Lu et al., JGR 2012]

#### Ne Variations during the course of PPEF



## Comparison of TIMEGCM and SAMI Ne



TIMEGCM

#### TIMEGCM O<sup>+</sup> Flux at 680 km



#### SAMI O<sup>+</sup> Flux at 680 km



## Evolution of HmF<sub>2</sub> During PPEF





## TEC Maps at Selected UT During PPEF



## TIMEGCM TEC (below ~680km)



#### Comparison of Observed and Modeled Ne



#### Comparison of Observed and Modeled TEC at 75°W



# Summary

- The coupled model eliminates the upper boundary limitations of the TIMEGCM, showing a much improved agreement with observations;
- During PPEF, the super-fountain effect can lift the F-layer above 800 km;
- The O<sup>+</sup> fluxes exhibits complex latitudinal and longitudinal variations during PPEF, with strong upward and downward flows in localized regions, making it very difficult to parameterize the O<sup>+</sup> flux in the top ionosphere.
- Future improvements are needed:
  - realistic magnetic fields (i.e., IGRF)
  - two-way coupling



[Lu et al., JGR 2012]



[Lu et al., JGR 2012]

#### SAMI TEC (below 680km) Maps



#### Ratio of TEC (below 680km)/TEC



## Two Major Limitations of the TIMEGCM

