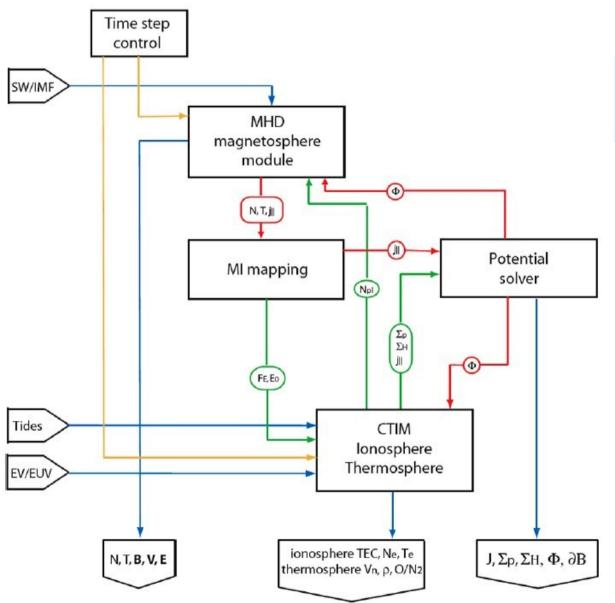
GEM 2014 Initial results of storm-time IT responses using OpenGGCM-CTIM global magnetosphereionosphere-thermosphere coupling model

H. K. Connor<sup>1</sup>, M. Fedrizzi<sup>2</sup>, Y. Shi<sup>1</sup>, E. Zesta<sup>1</sup>, T. Fuller-Rowell<sup>2</sup>, and J. Raeder<sup>3</sup> <sup>1</sup>NASA/GSFC, <sup>2</sup>NOAA, <sup>3</sup>UNH

**Purpose of this study** 

- 1. Test the OpenGGCM-CTIM coupling model
- 2. Understand the importance of magnetospheric energy input to predict storm-time IT system.

# Model 1: OpenGGCM-CTIM



Two-way coupled magnetosphereionosphere-thermosphere model.

> **Current conservation**  $\nabla \cdot \underline{\Sigma} \cdot \nabla \Phi = -J_{||} \sin I$

OpenGGCM calculates the magnetosphere part by solving resistive MHD equations with solar wind input.

CTIM calculates the ionosphere and thermosphere by solving both neutral and ion fluid equations self-Consistently.

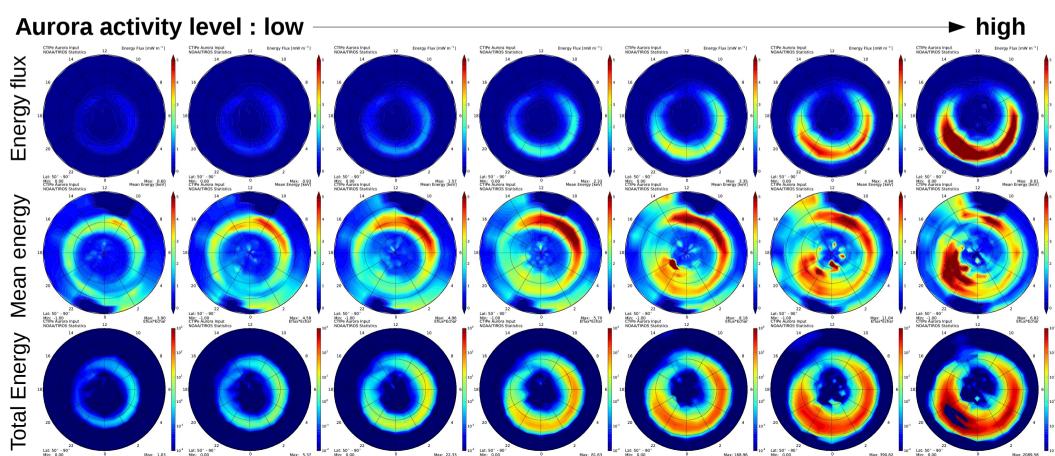
OpenGGCM coupled with CTIM provides realistic ionospheric potential patterns because the CTIM calculates the ionospheric conductance more accurately.

# Model 2: CTIPe

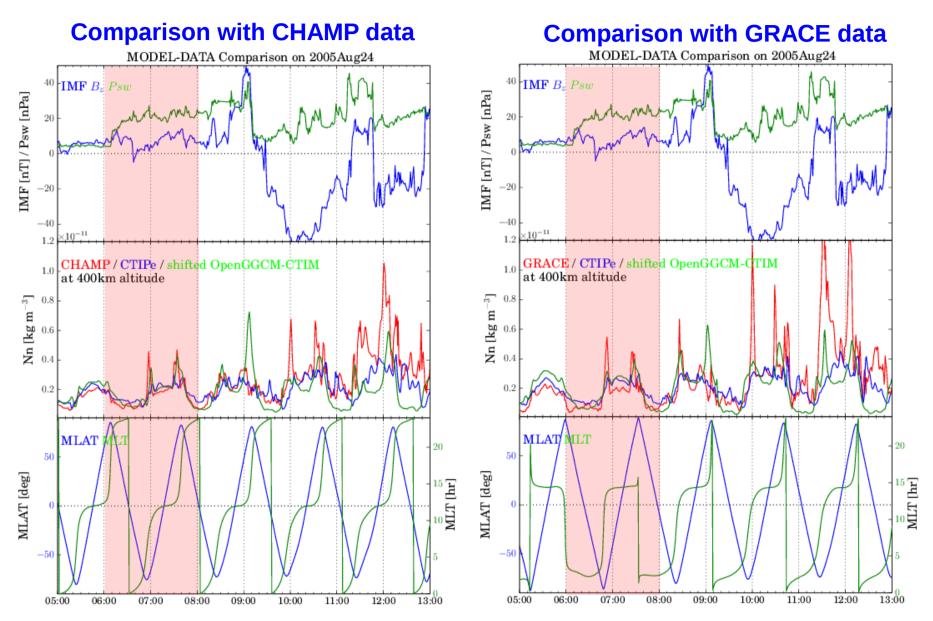
- The latest version of CTIM.
- Dynamics of a low- and mid- latitude ionosphere, a plasmasphere, and global dynamo electric fields are added to this version.
- Magnetospheric energy input:

Electric field from Weimer 2005

Auroral precipitation from NOAA/TIROS Statistics

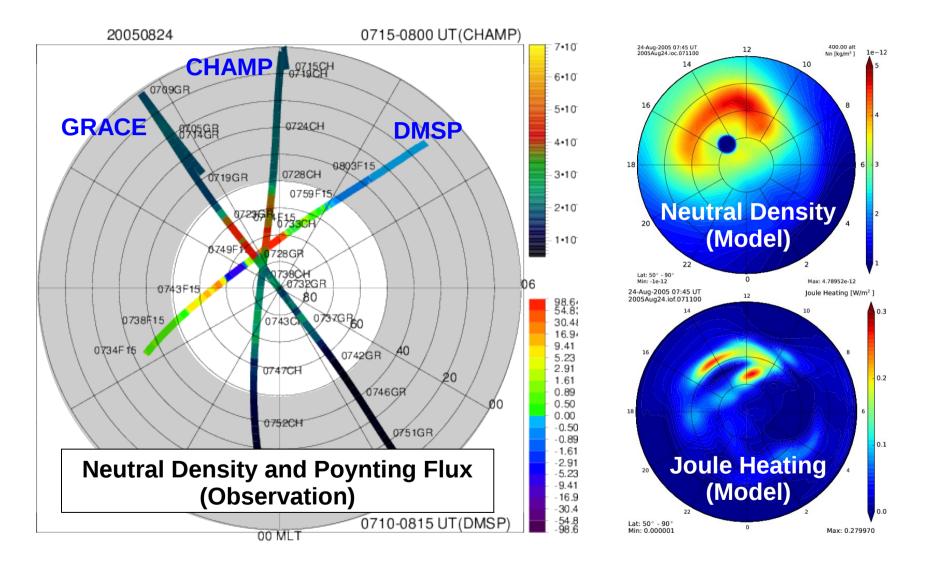


### 2005-08-24 CME STORM



OpenGGCM-CTIM shows enhancement of high-latitude neutral density during the strong Psw period and the early storm phase.

# **MODEL-DATA comparison**



Observations from CHAMP, GRACE, and DMSP match with the model calculation of neutral density and Joule heating.

## **Model-Model comparison of Psw impact**

**CTIPe** 

Electric potential

Aurora power

**Neutral Density** 

**OpenGGCM-CTIM** 

Potential [kV]

100

-20

-40

-60

-80

100

Max: 64 306000

eflux\*echar

Max: 531,277

Max: 3.84752e-12

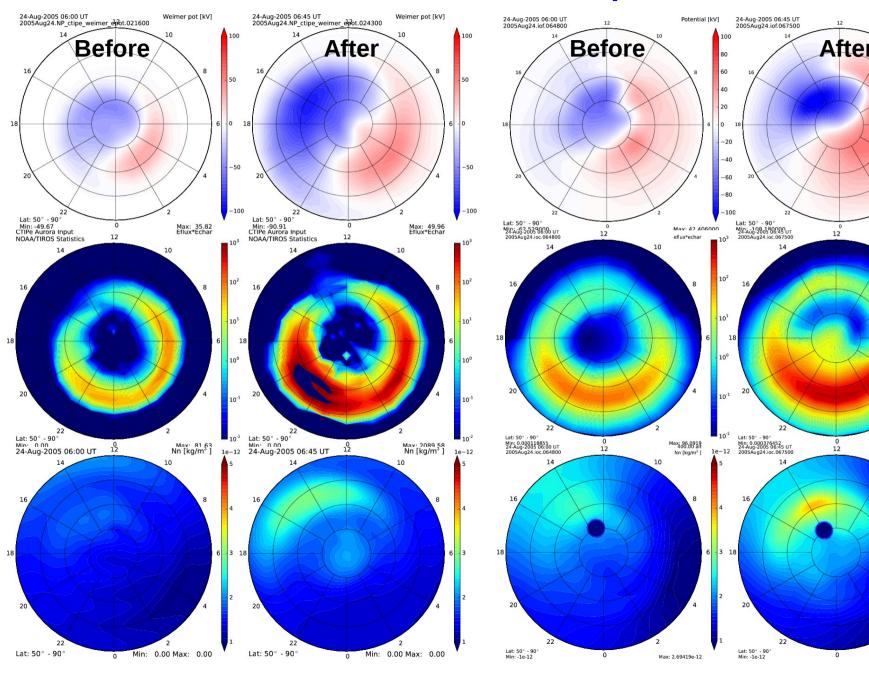
10

Nn [kg/m<sup>3</sup>]

1e-12

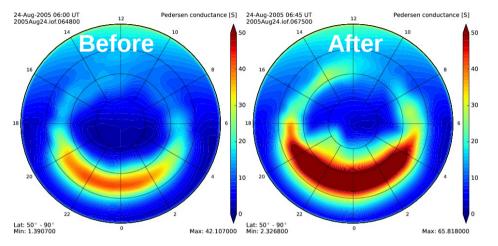
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12

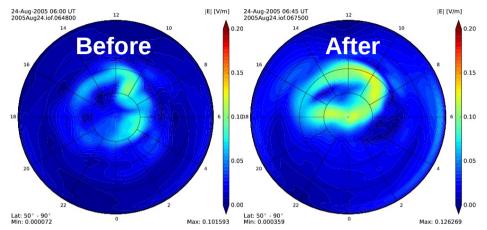


### **Thermosphere Heating after Psw impact**

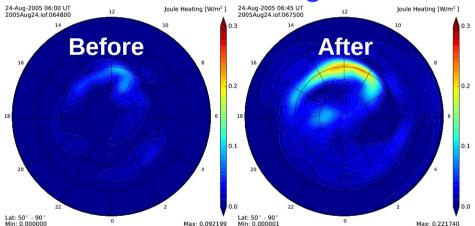
#### **Pederson Conductivity**



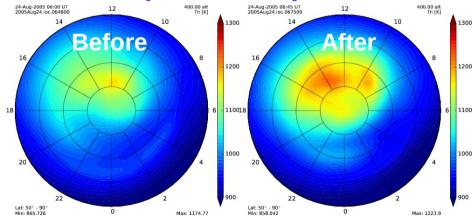
#### **Total Electric Field**



**Joule Heating** 



#### **Thermosphere Temperature**



Enhanced electric field and Pederson conductivity heat the dayside thermosphere, initiating upwelling of neutrals and increasing neutral density of this region.

## SUMMARY

- We simulate a geomagnetic storm on Aug 24, 2005 using OpenGGCM-CTIM in order to test storm-time IT responses.
- OpenGGCM-CTIM reproduces high-latitude neutral density peaks, showing a good agreement with CHAMP and GRACE observations.
- Both electric field and aurora precipitation increases during the compression period, heating thermosphere via joule heating and initiating neutral upwelling. This leads to neutral density enhancement at high latitude.

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# **OpenGGCM-CTIM**

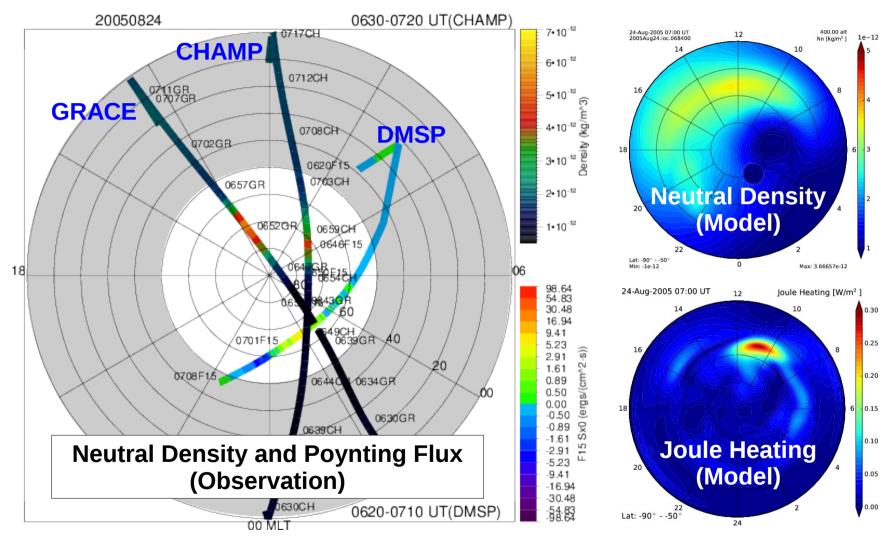
- Two-way coupled magnetosphere-ionosphere-thermosphere model.
- OpenGGCM calculates magnetosphere and planar ionosphere by solving resistive MHD equations with solar wind and IMF input.
- CTIM calculates global thermosphere and high-latitude ionosphere three-dimensionally by solving neutral and ion fluid equations.
- OpenGGCM-CTIM magnetospheric energy input
  - Electric potential from current conservation:

$$\nabla \cdot \underline{\Sigma} \cdot \nabla \Phi = -J_{||} \sin I$$

- Aurora precipitation

Diffuse aurora :  $F_E = n_e (kT_e/2\pi m_e)^{0.5}$ ,  $E_0 = kT_e$ Discrete aurora :  $F_E = \Delta \Phi_{||} J_{||}$ ,  $E_0 = e \Delta \Phi_{||}$ 

# **Model-DATA comparison**



OpenGGCM-CTIM and satellite observations show a good agreement of high neutral density regions.