

GEM-CEDAR Challenge: Comparing Ionospheric Models with Poynting Flux from DMSP Observations

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<http://ccmc.gsfc.nasa.gov>



Poynting Flux and Joule Dissipation

Poynting Flux:

$$PF = \left(\frac{ExB}{\mu_0} \right)_z$$

Obtained along DMSP satellite track from E and B measurements.
Mostly negative (inflow of electromagnetic energy).

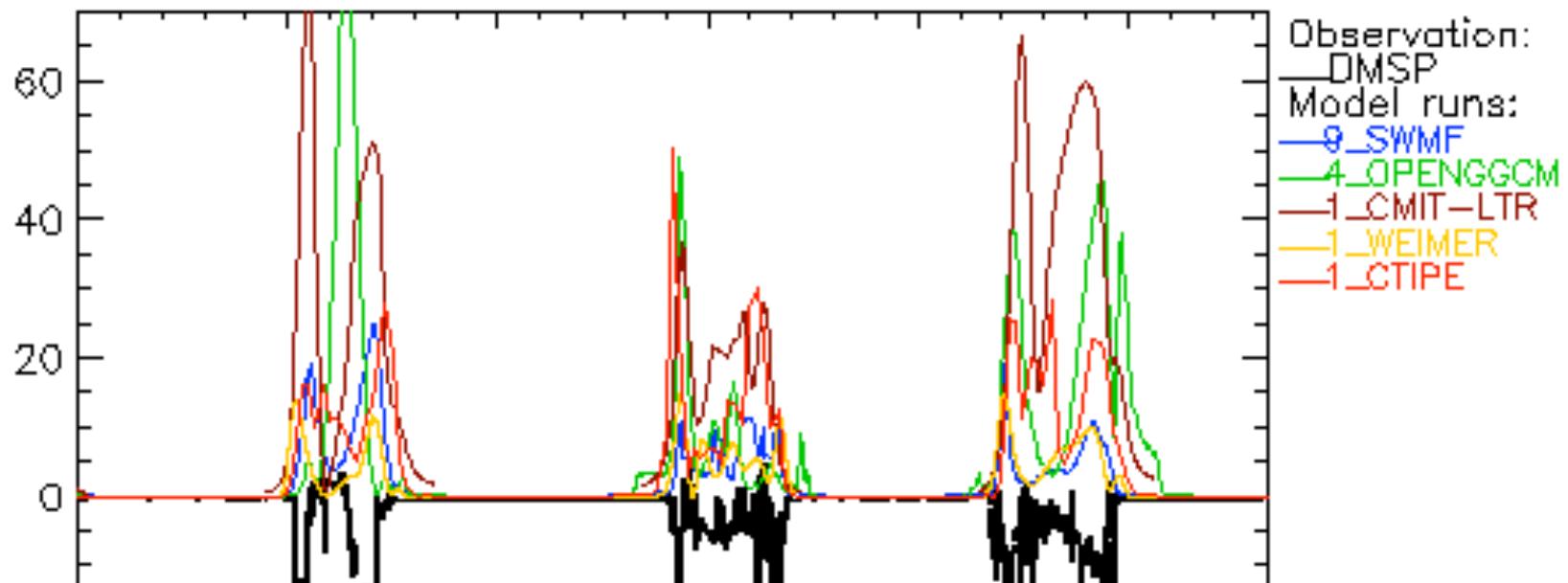
Joule Heating:

$$JH = \frac{J^2}{\Sigma_P}$$

Use height-integrated currents and Pedersen conductance.
Always positive (drain on electromagnetic energy).

DMSP orbits

Sz from observatory file: OBS_DMSP.txt



Model runs:

9_SWMF: 2011 SWPC model version; realtime with 1M cells,
SWPC_SWMF_052811_4

4_OPENGCCM: OpenGGCM V. 4.0 real-time setup 3.88M cells,
rotating dipole, SWPC_OpenGGCM_031311_4

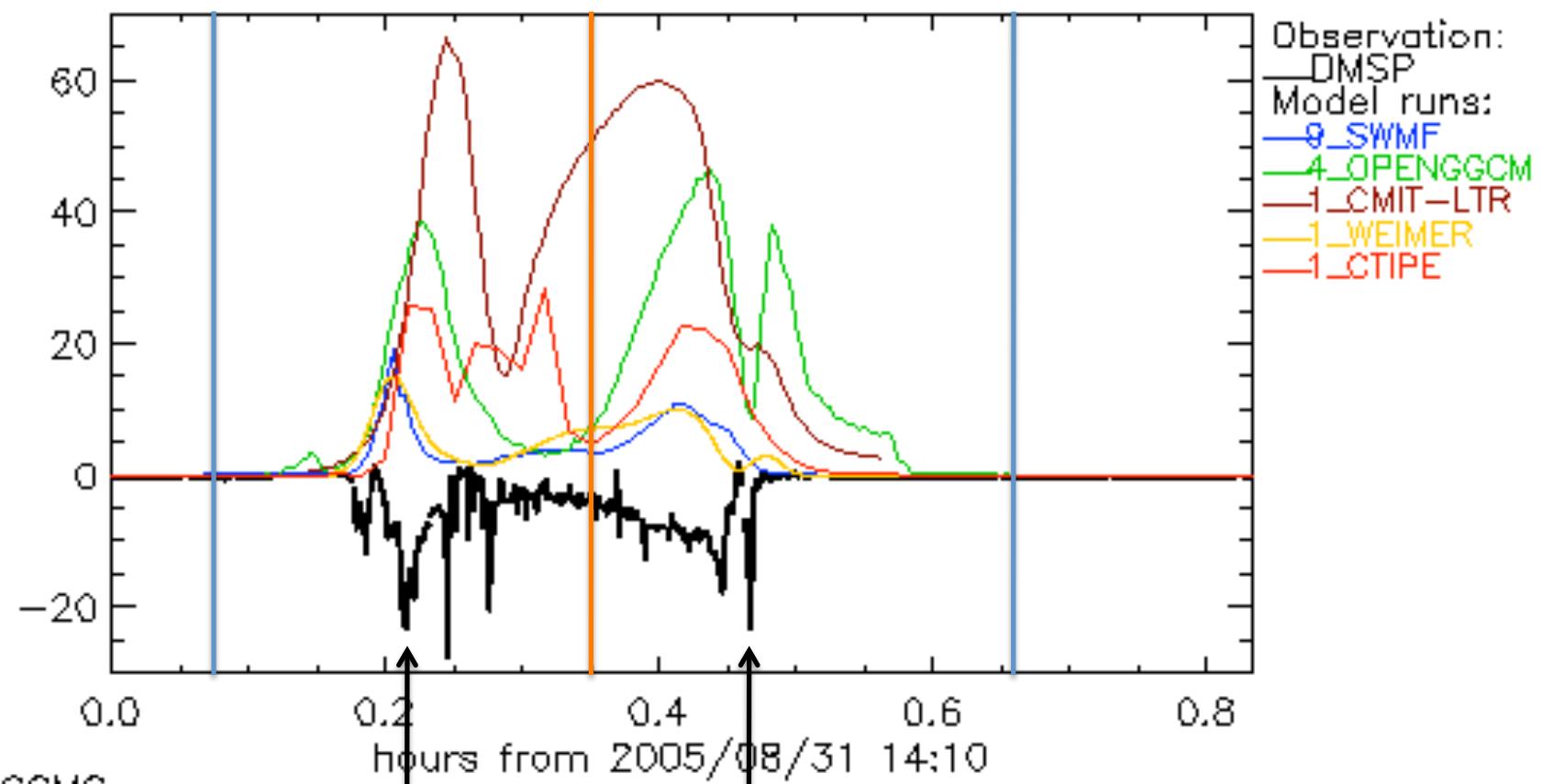
1_CMIT-LTR: 53x48x64 grid, real-time setup (SWPC_CMIT-LFM-MIX_031711_4)

1_WEIMER: results of Weimer-2005 (submitted Dec. 12, 2009)

1_CTIPE: CTIPE 2.0, run SWPC_CMIT-LFM-MIX_031711_4

Analysis of single pass

Sz from observatory file: OBS_DMSP.txt

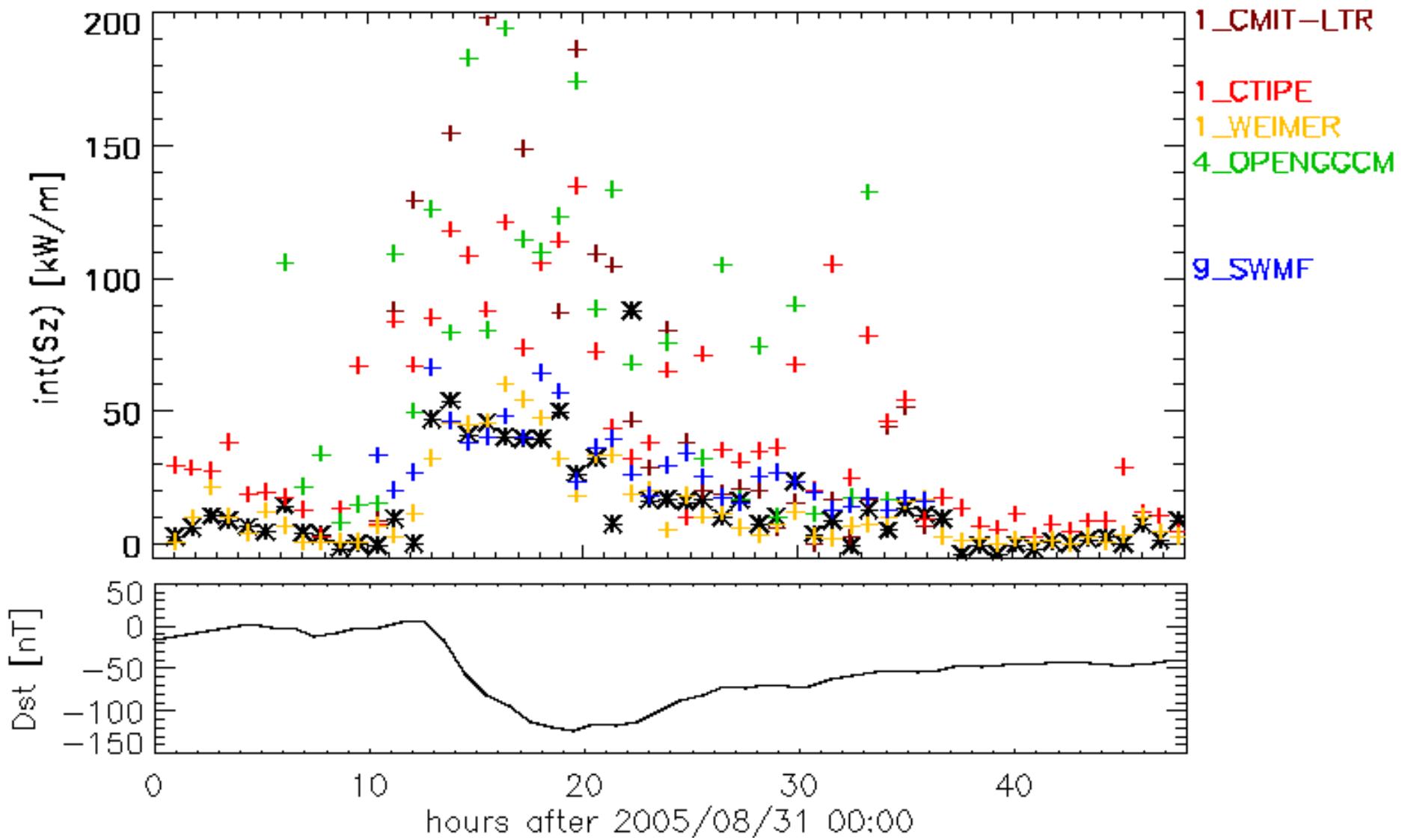


Plot: CCMC

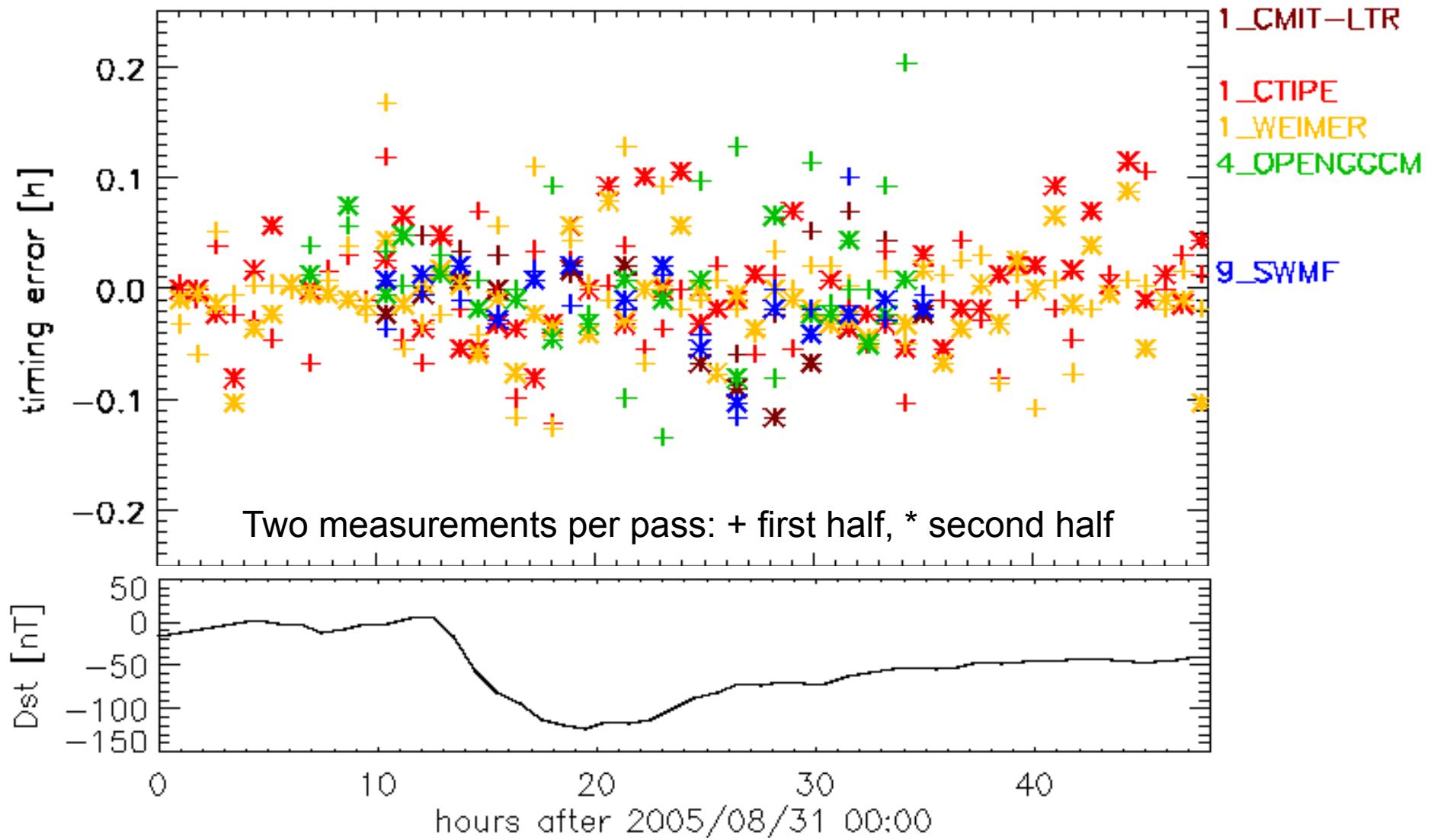
Observed maxima

en DMSP leaves auroral zone

Integrated Joule Heating vs. PF

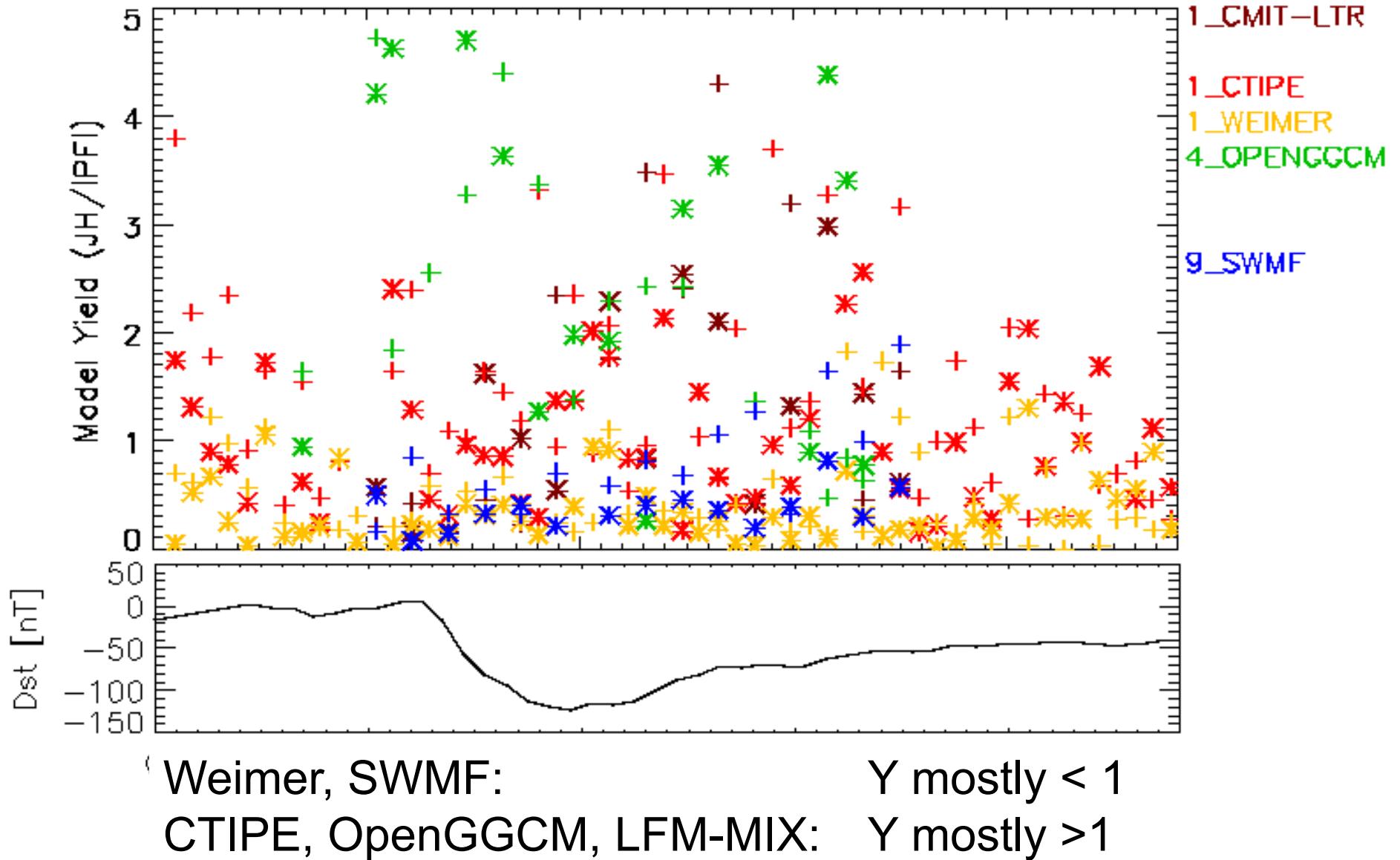


Timing errors (ideal: zero)



No systematic trend: errors go both ways for either segment of pass, not significant variation with activity level.

Model Yields (ideal score: 1)



Summary

- Included:
 - statistical model (Weimer)
 - first-principles ionosphere-thermosphere (CTIPe)
 - Electrodynamics of global magnetosphere models (SWMF, OpenGGCM, LFM-MIX)
- Needed:
 - TIE-GCM, USU-GAIM: electrodynamics
 - other coupled models (RCM-CTIPe, ...)
 - LFM-MIX-TIEGCM (can now run at CCMC)
 - Work on other events in challenge.